

Demystifying Wheeler-de Witt

Aidan Chatwin-Davies

July 5, 2016

1 Exercise: Minisuperspace

Consider the *minisuperspace* that consists of all closed homogeneous and isotropic FRW spacetimes in 1 + 3 dimensions that contain a homogeneous and isotropic real scalar field, Φ . For fun, let's also suppose that there is a (non-negative) cosmological constant Λ that is fixed to some value, *i.e.*, we don't consider it as part of the minisuperspace. In a suitable coordinate system, the line element of the spacetime manifold reads

$$ds^2 = \frac{3}{\Lambda} (-dt^2 + a^2(t)d\Omega_3^2), \quad (1)$$

where $d\Omega_3^2$ is the round metric on the 3-sphere. In this setting, a 3-geometry and matter configuration is entirely characterized by only two real numbers: the value of the scale factor, b , and the value of the scalar field, Φ_0 . As such, the wavefunctional over geometries and matter configurations reduces to a wavefunction with two arguments: $\Psi[\gamma_{ij}(\mathbf{x}), \{\Phi(\mathbf{x})\}] \equiv \Psi(b, \Phi_0)$. In this exercise, we will deduce the form of the Wheeler-De Witt equation that this wavefunction obeys.

(a) In units where $G = \hbar = c = 1$, the Einstein-Hilbert action is

$$S_{EH} = -\frac{1}{16\pi} \int d^4x \sqrt{-g} (R - 2\Lambda) \quad (2)$$

and the action of a free scalar field is

$$S_\Phi = \frac{1}{2} \int d^4x \sqrt{-g} (g^{\mu\nu} \Phi_{,\mu} \Phi_{,\nu} + m^2 \Phi^2). \quad (3)$$

Use the form of the metric above and the fact that $\Phi(t)$ depends only on t to write down the total action $S = S_{EH} + S_\Phi$. You may find it convenient to express your answer in terms of the following quantities:

$$H^2 \equiv \frac{\Lambda}{3} \quad \phi \equiv \sqrt{\frac{4\pi}{3}} \Phi \quad \mu^2 \equiv \frac{3}{\Lambda} m^2 \quad (4)$$

(b) We can read off the Lagrangian \mathcal{L} from the total action $S = \int dt \mathcal{L}$. Take a and ϕ to be the canonical coordinates and define their conjugate momenta via

$$p_a = \frac{\partial \mathcal{L}}{\partial \dot{a}} \quad p_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}}, \quad (5)$$

where a dot over a quantity denotes a derivative with respect to t . Perform a Legendre transformation and write down the Hamiltonian \mathcal{H} in terms of a , ϕ , p_a , and p_ϕ . Note that \mathcal{H} has no explicit dependence on the coordinate time t ; at any given time t_0 , \mathcal{H} depends only on the values of $a(t_0) \equiv b$, $\phi(t_0) \equiv \chi$, $p_a(t_0) \equiv p_b$, and $p_\phi(t_0) \equiv p_\chi$. As such, to emphasize that \mathcal{H} is purely a function of 3-geometry and the matter configuration, we can write $\mathcal{H} = \mathcal{H}(b, \chi, p_b, p_\chi)$.

(c) Note that there do not seem to be any secondary constraints for the momenta above, *i.e.*, $\dot{p}_a, \dot{p}_\phi \neq 0$. Why is this? Is it still correct to write $\mathcal{H} = 0$?

(d) Make the identification

$$p_b \rightarrow -i \frac{\partial}{\partial b} \quad p_\chi \rightarrow -i \frac{\partial}{\partial \chi} \quad (6)$$

and write down the explicit form of the Wheeler-de Witt equation $\mathcal{H}\Psi(b, \chi) = 0$.

(e) Express the Wheeler-De Witt operator $\mathcal{H}(b, \chi, \partial_b, \partial_\chi)$ in the following form:

$$\mathcal{H} = \frac{1}{2} G^{AB} \partial_A \partial_B + \mathcal{V} \quad (7)$$

G_{AB} and \mathcal{V} are the Wheeler-De Witt metric and potential for minisuperspace in the coordinates $x^A = (b, \chi)$. What is the signature of the Wheeler-De Witt metric?